Methyl mercaptan (CH₃SH) as a probe for variation of the proton-to-electron mass ratio

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Torsion-rotation transitions in molecules exhibiting hindered internal rotation possess enhanced sensitivities to a variation of the proton-to-electron mass ratio. This enhancement occurs due to a cancellation of energies associated with the torsional and rotational degrees of freedom of the molecule. This effect occurs generally in every internal rotor molecule, but is exceptionally large in methanol. In this paper, we calculate the sensitivity coefficients of methyl mercaptan, which is the thiol analog of methanol. The obtained sensitivity coefficients in this molecule range from $K_{\mu} = -14.8$ to +12.2 for transitions with a lower-level excitation energy below 10 cm^{-1} .

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I. INTRODUCTION

Physical theories extending the standard model of particle physics have presented scenarios that allow for spatialtemporal variations of the constants of nature [1]. Since the initial findings of a possible variation of the fine-structure constant by Webb et al. [2], there has arisen a great activity in the search for signatures of such variations. Studies aimed at detecting a possible drift of a fundamental constant on a cosmological time scale focus mainly on the fine-structure constant α [3–6] and the proton-to-electron mass ratio μ [7–9]. A variation of α or μ will manifest itself as a change in the spectrum of atoms and molecules, since such a variation may induce a shift in the position of a spectral line. Not all lines will shift in the same amount or direction. The response of a transition to a variation of α or μ is characterized by its sensitivity coefficient, K_{μ} or K_{α} , respectively, which is defined as the proportionality constant between the fractional frequency shift of the transition, $\Delta v / v$, and the fractional shift in α or μ ,

$$\frac{\Delta \nu}{\nu} = K_X \frac{\Delta X}{X}$$
 with $X = \alpha, \mu.$ (1)

The search for μ variation on a cosmological time scale has been made operational by comparing optical transitions of molecular hydrogen (H₂) in high-redshifted objects with accurate laboratory measurements [7]. These investigations have yielded a limit at the level of $\Delta \mu / \mu < 10^{-5}$ for look-back times of 12 billion years [8,9]. The transitions in H_2 that were used to obtain this result possess sensitivity coefficients that range from -0.05 to +0.02. Inversion transitions of ammonia (NH₃) were found to be ~ 100 times more sensitive to μ variation than H₂ transitions [10,11]. Astronomical observations of NH₃, in the microwave or radio range of the electromagnetic spectrum, led to stringent 1σ constraints at the level of $(1.0 \pm 4.7) \times 10^{-7}$ [12] and $(3.5 \pm 1.2) \times 10^{-7}$ [13]. Soon thereafter, it was realized that the large number of degrees of freedom that exist in even the simplest polyatomic molecules can result in large enhancements of the sensitivity coefficients for a possible drift in μ . These enhancements

occur for transitions between near-degenerate levels that each have a different dependence on μ . For instance, it was found that mixed inversion-rotation transitions in H₃O⁺ have sensitivity coefficients ranging from $K_{\mu} = -9$ to +5.7 [14], while the Renner-Teller interaction in *l*-CH₃ results in sensitivity coefficients ranging from $K_{\mu} = -53$ to +742 [15]. Mixed torsion-wagging-rotation transitions in methylamine display sensitivity coefficients ranging from $K_{\mu} = -19.1$ to -0.75 [16].

In the context of astrophysical searches, methanol [17,18] is the target species of choice since it possesses sensitive transitions at low excitation energy and has been observed at high redshift [19]. In a recent study, Bagdonaite *et al.* used four transitions in methanol to constrain $\Delta \mu/\mu$ at $(0.0 \pm 1.0) \times 10^{-7}$ at a look-back time of 7 billion years [20]. Methyl mercaptan (CH₃SH) is the sulfur analog of methanol and might therefore possess transitions that have large sensitivity coefficients to a variation of μ [21]. Although methyl mercaptan has thus far only been detected in our local galaxy [22,23], recent advances in radio telescopes have greatly increased the number of detected molecular species at high redshift. It is therefore relevant to have a list available that contains the sensitivity coefficients of transitions in methyl mercaptan that might be observed in the interstellar medium.

The recent terahertz and far-infrared study of the normal isotopologue (${}^{12}CH_{3}{}^{32}SH$) of methyl mercaptan by Xu *et al.* [24] has resulted in a complete list of the molecular parameters for this molecule. In this paper, we use the results of Xu *et al.* [24] and the scaling relations of the molecular parameters that were derived in Jansen *et al.* [17,21] to calculate the sensitivity coefficients of methyl mercaptan.

II. STRUCTURE OF METHYL MERCAPTAN

Methyl mercaptan or methanethiol, depicted on the righthand side of Fig. 1, consists of a thiol group (SH) attached to a methyl group (CH₃) and is thus the sulfur analog of methanol. The CS σ bond connecting the two parts of the molecule is flexible, allowing the methyl group to rotate with respect to the thiol group. As in the case of



FIG. 1. (Color online) Variation of the potential energy of methyl mercaptan with the relative rotation γ of the SH group with respect to the methyl group about the molecular axis. Shown are the J = 1, |K| = 1 energies of the lowest torsion-vibrational levels. The splitting between the different symmetry levels is due to tunneling through the potential barriers. The *A*-symmetry species are split further due to the asymmetry of the molecule (*K* splitting).

methanol, this rotation is not free but hindered by a threefold potential barrier with minima and maxima that correspond to the staggered and eclipsed configuration of the molecule, respectively. For the lowest-energy levels, this relative or internal rotation is classically forbidden and only occurs due to quantum-mechanical tunneling of the hydrogen atoms. As a consequence of this tunneling, each rotational level splits into three levels that are labeled according to their symmetry as Aor E, as can be seen on the left-hand side of Fig. 1.

The lowest-energy levels of CH_3SH are shown in the left and right panel of Fig. 2. The A and E species can be considered as two different molecular species in the same sense as ortho- and para-ammonia, respectively. The arrangement of energy levels within a symmetry state resembles that of a prolate symmetric top, with the difference being that every Kladder obtains a small energy offset due to the K-dependent tunneling splitting.

As a consequence, certain states in neighboring K ladders may become near degenerate, which results in a large enhancement of the sensitivity coefficients K_{μ} for transitions connecting these states [17,18].

III. SENSITIVITY COEFFICIENTS

The energy levels of methyl mercaptan have been calculated using an adapted version of the BELGI code [25]. This code was modified and improved by Xu *et al.* [26] in a number of ways useful for treating the large data sets available for the methanol molecule, but the general approach has not been changed compared to the initial code. The present version of the code has been sped up compared to the original version, and also a substantial number of higher-order parameters has been added. Using the set of 79 molecular parameters of methyl mercaptan obtained by Xu *et al.*, the lower-energy levels are found with an accuracy of <100 kHz.

In order to calculate the K_{μ} coefficients for the different transitions in methyl mercaptan, the energy of each level and its dependence on μ has to be obtained. This translates into knowing the values of the molecular constants that go into BELGI and how these constants scale with μ . These scaling relations were obtained in a previous study [17], while a more detailed discussion on the derivation can be found in Ref. [21].

Table I lists calculated transition frequencies and sensitivity coefficients in methyl mercaptan with an excitation energy



FIG. 2. Energy of the lowest rotational levels in the torsion-vibrational ground state ($v_t = 0$) of methyl mercaptan (${}^{12}CH_3{}^{32}SH$) [24]. The levels are denoted by J_K (indicated on the left side of each level). For the *A* levels, the so-called parity quantum number (+ or -) is also used. The panel on the left displays the *A* state levels, whereas the panel on the right displays the *E* state levels. High sensitivities are expected for transitions that connect near-degenerate levels with different *K*.

TABLE I. Transition frequencies and sensitivity coefficients in methyl mercaptan with a lower-level excitation energy of less than 10 cm⁻¹, calculated with BELGI using the molecular constants of Ref. [24] (fifth column) and the toy model of Ref. [21] (sixth column). The third column lists the transition strength multiplied by the electric dipole moment, μ_e , squared, while the fourth column lists the excitation energy of the lower level in Kelvin.

Transition	Energy (MHz)	$S\mu_e^2({\rm D}^2)$	$T_{\rm low}~({\rm K})$	$K_{\mu}^{ m BELGI}$	K_{μ}^{toy}
$1_1 \rightarrow 1_1 A^{\pm}$	523.147	1.219	6.3	-0.98	-1.00
$2_1 \rightarrow 2_1 A^{\pm}$	1569.410	0.677	8.7	-0.98	-1.00
$3_1 \rightarrow 4_0 E$	1874.635	1.390	12.0	12.20	11.77
$4_0 \rightarrow 3_1 A^+$	3038.566	3.077	12.1	-14.83	-14.94
$3_1 \rightarrow 3_1 A^{\pm}$	3138.723	0.473	12.3	-0.98	-1.00
$2_{-1} \rightarrow 3_0 E$	10534.181	1.064	6.8	-7.55	-7.29
$2_0 \rightarrow 1_{-1}E$	14764.687	0.513	3.6	3.68	3.49
$3_0 \rightarrow 2_1 E$	23339.083	0.966	7.3	-2.07	-2.03
$0_0 \rightarrow 1_0 A^+$	25290.869	0.813	0.0	-1.00	-1.00
$0_0 \rightarrow 1_0 E$	25291.824	0.814	0.0	-1.00	-1.00
$3_0 \rightarrow 2_1 A^+$	29091.802	2.038	7.3	-2.44	-2.46
$3_{-1} \rightarrow 4_0 E$	35857.370	1.678	10.4	-2.92	-2.85
$2_0 \rightarrow 1_1 E$	48604.208	0.496	3.6	-1.51	-1.49
$1_1 \rightarrow 2_1 A^+$	50058.794	1.220	6.3	-1.00	-1.00
$1_{-1} \rightarrow 2_{-1}E$	50565.538	1.220	4.3	-1.00	-1.00
$1_0 \rightarrow 2_0 A^+$	50579.301	1.625	1.2	-1.00	-1.00
$1_0 \rightarrow 2_0 E$	50580.882	1.629	1.2	-1.00	-1.00
$1_1 \rightarrow 2_1 E$	50599.280	1.221	6.0	-1.00	-1.00
$1_1 \rightarrow 2_1 A^-$	51105.057	1.220	6.3	-1.00	-1.00
$2_0 \rightarrow 1_1 A^+$	54895.867	1.014	3.6	-1.76	-1.77
$4_0 \rightarrow 4_{-1}E$	65172.338	3.774	12.1	0.05	0.02
$3_0 \rightarrow 3_{-1}E$	65282.263	3.162	7.3	0.06	0.02
$2_0 \rightarrow 2_1 E$	65330.225	2.383	3.6	0.06	0.01
$1_0 \rightarrow 1_1 E$	65345.568	1.480	1.2	0.06	0.01
$2_1 \rightarrow 3_1 A^+$	75085.877	2.168	8.7	-1.00	-1.00
$2 \rightarrow 3 \rightarrow E$	75816.443	2.168	6.8	-1.00	-1.00
$2_{0} \rightarrow 3_{0}A^{+}$	75862.860	2.438	3.6	-1.00	-1.00
$2_0 \rightarrow 3_0 E$	75864.406	2.443	3.6	-1.00	-1.00
$2_0 \rightarrow 3_1 E$	75925.915	2.169	8.4	-1.00	-1.00
$2_1 \rightarrow 3_1 A^-$	76655 189	2.168	8.8	-1.00	-1.00
$0_0 \rightarrow 1_1 E$	90637 393	1 017	0.0	-0.24	-0.27
$1_0 \rightarrow 1_1 E$	99185 090	1.520	1.2	-1.25	-1.24
$2_0 \rightarrow 2_1 E$	99203 488	2.602	3.6	-1.25	-1.21
$2_0 \rightarrow 2_1 E$ $3_0 \rightarrow 3_1 E$	99264 998	3 784	73	-1.25	-1.21
$4_0 \rightarrow 4_1 F$	99409 714	5 101	12.1	-1.24	_1.24
$3_1 \rightarrow 4_1 A^+$	100110 190	3 049	12.1	-1.00	-1.00
$3 \rightarrow 4 \rightarrow E$	101029 708	3.048	10.4	-1.00	-1.00
$3_0 \rightarrow 4_0 A^+$	101139.112	3.251	7.3	-1.00	-1.00
$3_0 \rightarrow 4_0 E$	101139 633	3 257	73	-1.00	-1.00
$3_1 \rightarrow 4_1 E$	101284 349	3 049	12.0	-1.00	-1.00
$3_1 \rightarrow 4_1 A^-$	102202 438	3 049	12.0	-1.00	-1.00
$1_0 \rightarrow 1_1 A^{\pm}$	105998 315	3.018	12.1	-1.00	-140
$2_0 \rightarrow 2_1 A^{\pm}$	106524 072	5.017	3.6	-1.39	-140
$2_0 \rightarrow 2_1 M$ $3_0 \rightarrow 3_1 A^{\pm}$	107316 401	6 996	73	-1.39	-1.39
$4_0 \rightarrow 4_1 A^{\pm}$	108379 727	8 947	12.1	-1.39	-1.39
$1_0 \rightarrow 2_1 F$	115911 107	1 572	12.1	-0.40	-0.43
$0_0 \rightarrow 1_1 F$	124476 914	0.983	0.0	-1.20	-1 19
$0_0 \rightarrow 1_1 E$	124476 914	0.983	0.0	-1.20	_1 10
$4_0 \rightarrow 5_0 F$	126403 807	4 071	12.1	-1.00	-1.00
$4_0 \rightarrow 5_0 A^+$	126405 629	4 063	12.1	-1.00	-1.00
$0_0 \rightarrow 1.4^+$	130766 037	2.012	0.0	_1.00	_1 32
$2_0 \rightarrow 3_{-}F$	141146 668	2.012	3.6	-0.51	-0.53
$\frac{1}{1_0} \rightarrow 2_1 E$	149784.370	1.428	1.2	-1.17	-1.16

TABLE I. (Continued.)

Transition	Energy (MHz)	$S\mu_e^2({ m D}^2)$	$T_{\rm low}$ (K)	$K_{\mu}^{ m BELGI}$	K_{μ}^{toy}
$1_0 \rightarrow 2_1 A^+$	155533.962	3.018	1.2	-1.27	-1.27
$3_0 \rightarrow 4_{-1}E$	166311.971	2.896	7.3	-0.59	-0.60
$3_1 \rightarrow 2_2 E$	172960.517	0.333	12.0	-0.64	-0.67
$2_0 \rightarrow 3_1 E$	175129.404	1.816	3.6	-1.14	-1.14
$2_0 \rightarrow 3_1 A^+$	180040.538	4.029	3.6	-1.23	-1.24
$4_0 \rightarrow 5_{-1}E$	191366.909	3.710	12.1	-0.64	-0.65
$3_{-1} \rightarrow 2_{-2}E$	193330.838	0.335	10.4	-0.97	-0.98
$3_0 \rightarrow 4_1 E$	200549.347	2.127	7.3	-1.12	-1.12
$3_0 \rightarrow 4_1 A^+$	204287.868	5.050	7.3	-1.21	-1.21
$3_1 \rightarrow 2_2 A^-$	215287.322	0.337	12.4	-1.28	-1.28
$3_1 \rightarrow 2_2 A^+$	218428.095	0.332	12.3	-1.27	-1.27
$4_0 \rightarrow 5_1 E$	226093.110	2.345	12.1	-1.10	-1.11
$4_0 \rightarrow 5_1 A^+$	228279.583	6.084	12.1	-1.18	-1.19
$3_1 \rightarrow 3_2 E$	248835.455	2.929	12.0	-0.75	-0.77
$2_1 \rightarrow 2_2 E$	248886.432	1.672	8.4	-0.75	-0.77
$2_{-1} \rightarrow 2_{-2}E$	269147.282	1.684	6.8	-0.98	-0.99
$3_{-1} \rightarrow 3_{-2}E$	269204.583	2.949	10.4	-0.98	-0.99
$3_1 \rightarrow 3_2 A^{\mp}$	291169.820	2.957	12.4	-1.20	-1.20
$2_1 \rightarrow 2_2 A^{\mp}$	291944.561	1.680	8.8	-1.20	-1.20
$2_1 \rightarrow 2_2 A^{\pm}$	293511.922	1.667	8.7	-1.20	-1.20
$3_1 \rightarrow 3_2 A^{\pm}$	294298.297	2.911	12.3	-1.20	-1.20
$1_1 \rightarrow 2_2 E$	299485.712	2.994	6.0	-0.79	-0.81
$1_{-1} \rightarrow 2_{-2}E$	319712.820	3.015	4.3	-0.98	-0.99
$2_1 \rightarrow 3_2 E$	324761.370	3.315	8.4	-0.81	-0.83
$1_1 \rightarrow 2_2 A^-$	343047.569	3.001	6.3	-1.17	-1.17
$1_1 \rightarrow 2_2 A^+$	343572.765	2.993	6.3	-1.17	-1.17
$2_{-1} \rightarrow 3_{-2}E$	345021.026	3.337	6.8	-0.98	-0.99
$3_1 \rightarrow 4_2 E$	350003.733	3.711	12.0	-0.82	-0.84
$2_1 \rightarrow 3_2 A^-$	367814.764	3.334	8.8	-1.16	-1.16
$2_1 \rightarrow 3_2 A^+$	369394.420	3.308	8.7	-1.16	-1.16
$3_{-1} \rightarrow 4_{-2}E$	370371.709	3.732	10.4	-0.98	-0.99
$3_1 \rightarrow 4_2 A^-$	392318.890	3.754	12.4	-1.15	-1.15
$3_1 \rightarrow 4_2 A^+$	395488.347	3.695	12.3	-1.15	-1.15

less than 10 cm^{-1} , that is, those transitions most relevant for astrophysical searches. It can be seen that several of these transitions display relatively large sensitivity coefficients.

In the last column of Table I, the results of the approximate model of Ref. [21] are listed. This "toy" model is derived for molecules that exhibit hindered internal rotation and contain a C_{3v} symmetry group. The model decomposes the energy of the molecule into a pure rotational and a pure torsional part. The rotational part is approximated by the well-known expression for the rotational energy levels of a slightly asymmetric top,

$$E_{\rm rot}(J,K) = \frac{1}{2}(B+C)J(J+1) + \left(A - \frac{B+C}{2}\right)K^2, \quad (2)$$

with $A = 3.428 \text{ cm}^{-1}$, $B = 0.432 \text{ cm}^{-1}$, and $C = 0.413 \text{ cm}^{-1}$ the rotational constants along the *a*, *b*, and *c* axis of the molecule, respectively. The torsional energy contribution is approximated by a Fourier expansion as

$$E_{\text{tors}}(K) = F\left[a_0 + a_1 \cos\left\{\frac{2\pi}{3}(\rho K + \sigma)\right\}\right], \qquad (3)$$

with $F = 15.040 \text{ cm}^{-1}$ the constant of the internal rotation, $\rho = 0.652$ a dimensionless constant reflecting the coupling between internal and overall rotation, and $\sigma = 0, \pm 1$ a



FIG. 3. (Color online) Torsional energies in the ground torsional state ($v_t = 0$) of methyl mercaptan obtained with BELGI for A (solid circles), E1 (open circles), and E2 (open triangles) levels as a function of K. The solid, dashed, and dotted curves are fits to Eq. (3) for A, E1, and E2 states. Note that only integer values of K have physical meaning.

constant relating to the torsional symmetry. The expansion coefficients a_0 and a_1 depend on the shape of the torsional potential. Since we are mainly interested in the torsional energy difference a_0 cancels, a_1 follows from

$$a_1 = A_1 s^{B_1} e^{-C_1 \sqrt{s}}, \tag{4}$$

with $A_1 = -5.296$, $B_1 = 1.111$, and $C_1 = 2.120$ [21]. The dimensionless parameter $s = 4V_3/9F$, with $V_3 = 441.442$ cm⁻¹ the height of the barrier, is a measure of the effective potential. The torsional energy for methyl mercaptan is plotted in Fig. 3 as a function of *K*. Note that the torsional splitting between the *A* and *E* levels in the K = 0 state of methyl mercaptan is 0.99 cm⁻¹, and thus an order of magnitude smaller than in methanol, which has a torsional splitting of 9.1 cm⁻¹. As a consequence, the amount of energy that can be canceled in methyl mercaptan will be less than in methanol.

Finally, the sensitivity coefficient of the transition is obtained from

$$K_{\mu}^{\text{toy}} = \frac{K_{\mu}^{\text{rot}} \Delta E_{\text{rot}} + K_{\mu}^{\text{tors}} \Delta E_{\text{tors}}}{\Delta E_{\text{rot}} + \Delta E_{\text{tors}}}.$$
 (5)

Note that rather than using $\Delta E_{\text{rot}} + \Delta E_{\text{tors}}$, we chose to use the experimental energy difference between the levels, hv, in order to account for the slight asymmetry of the molecule. The sensitivity of a pure rotational transition is $K_{\mu}^{\text{rot}} = -1$, whereas the sensitivity of a pure torsional transition is given by $K_{\mu}^{\text{tors}} = (B_1 - 1) - \frac{1}{2}C_1\sqrt{s} = -3.7$ [21]. The sensitivity coefficients of this simple model are seen to agree well with the results obtained by a diagonalization of the full molecular Hamiltonian, which reflects the robustness of the obtained results.

IV. CONCLUSIONS

In this paper, we have calculated sensitivity coefficients for transitions between low-lying rotation levels in methyl mercaptan. The reported sensitivities span a range from $K_{\mu} =$ -14.8 to +12.2 and can therefore be used to search for the variation of μ in methyl mercaptan only. Although thus far methyl mercaptan has only been detected in our local galaxy, it is our hope that the advanced spectral coverage, resolution, and sensitivity of the new generation of radio telescopes such as ALMA (Atacama Large Millimeter Array) will result in the detection of this molecule at high redshift. Note that the comprehensive line list of accurate rest frequencies for methyl mercaptan obtained by Xu *et al.* [24] should also alleviate this search.

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